# Measurement of all-wave and spectral albedos of snow-covered summer sea ice in the Ross Sea, Antarctica 

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#### Abstract

All-wave albedo and spectral albedo data were collected over snowcovered pack-ice flocs during summer 1999 January and lebruary) in the Ross Sca, Antarctica. Temporal variation of the all-wave albedo and spectral albedo was measured from the northern edge to the southern edge of the pack icc along three lines of longitude: $165^{\circ} \mathrm{W}, 150^{\circ} \mathrm{W}$ and $135^{\circ} \mathrm{W}$. Snow depth, snow-cover stratification, snow-temperature profiles, and grain-size and morphology were also documented. It was observed that daily-averaged albedos were $0.70-0.86$ for cloudy conditions over the pack ice. Only two sets of daily-averaged albedo were collected for clear-sky conditions ( 0.788 and 0.825 ). Albedo was lower at the marginal edges of the pack ice than in the central pack ice. Albedo was higher over the southern part of the central pack ice than over the northern part. Clear- and cloudy-sky albedos measured on the same site indicate that the average increase in albedo duc to clouds is $1.4 \%$ (maximum $4.0 \%$. Spectral albedos of the packice floes were similar under clear-sky and cloudy conditions. Both are mainly controlled by the snow grain-size, especially in the top snow layer. All-wave albedos derived from measured visible and near-infrared spectral albedos, with extrapolation to ultraviolet and shortwave infrared regions for both clear- and cloudy-sky conditions, agreed well with all-wave albedos from direct measurements.


## INTRODUCTION

Spectral albedo is the ratio of the upwelling spectral irradiance reflected by the surface of snow and sea ice to the downwelling incident spectral irradiance at the surface. All-wave albedo is the integral of spectral albedo over the whole solar spectrum. It is a major variable controlling the energy budget of sca ice, and an important tuning parameter in modeling sea-ice-atmosphere interaction and global warming. Any reduction of the surface albedo means more absorption of solar radiation, which tends to increase the temperature at the surface. Absorption of shortwave radiation can result in internal melting and an increase of liquid-water content, changing the surface characteristics and internal structure of the snow cover Grenfell and Maykut, 1977). Accompanying the absorption of solar radiation is metamorphosis of snow crystals, which results in larger grain-si\%e and a decrease in number density of the snow grains. All of these factors will contribute to the decrease of snow albedo.

The all-wave albedo ofopen occan is much smaller than that of the snow and sea-ice cover. Approximately $8.5-95 \%$ of the incident solar radiation is absorbed by the open ocean, while $30-50 \%$ or less is absorbed by a snow and sea-ice cover (Washington and Parkinson, 1986). Thus, during the summer season, a reduction of snow and sea-ice albedo due to snow metamorphosis and possible removal of the snow cover entances the absorption of solar energy. This results in thimning and breaking-down of ice flocs. and possible reduction of ice concentration and icce extent, which will further reduce the area-averaged albedo. On the other hand, an increase in surface albedo due to freczing means more solar energy is reflected back to the atmosphere,
resulting in less imput of energy into the sea ice or ocean. This results in a higher probability of more freezing of ocean water in polar regions, and an increase in the areal ice coverage and possibly snow and sea-ice thickness, which will produce a further increase in the arca-averaged albedo. The positive feedback of the surface albedo to changes in atmospheric conditions makes the snow and sea ice in polar regions an important subject for weather prediction and global-warming studics.

In summer, the Arctic and the Antarctic have different patterns of sea-ice and snow decay (Andreas and Ackley, 1982; Fetterer and Unterstcincr, 1998). In the Aretic, melt ponds form as snow melts and meltwater accumulates on the ice surface. Decay of snow and ice is mainly from the top of the cover. In the Antarctic, melt ponds are rarely observed Arctowski, 1908; Wordie, 1921; Spichkin, 1967; Acklcy, 1979, so the surface albedo of Antarctic sca ice is close to the typical albedo valuc of snow-covered ice all year around. Antarctic sea-ice albedo shows large variation duc to the spatial complexity and variability of the surface even within the same region.

Alhough many albedo studies have been carried out in the Arctic Hanson, 1961; Chernigovski, 1966; Langleben, 1969, 1971: Grenfell and Maykul, 1977; Grenfell and Perovich, 1984; De Abreu and others, 1995) and Antarctic (Liljequist, 1956; Hanson, 1960; Hoinkes, 1960; Weller, 1968; Kuhn and others, 1977; Carroll and Fïtch, 1981; Yamanouchi, 1983; Allison and others. 1993; Grenfell and others, 1994), the data obtained are not sufficient for a global study. The processes affecting the snow surface albedo are still only qualitatively known, and the available albedo data on regional and temporal differences are too sparse to validate albedomodeling efforts (Fetterer and Lintersteiner, 1998). Most of
the albedo measurements in Antarctica are from continental sites. There are few datasets of all-wave and spectral albedos available for the Southern Ocean (e.g. Allison and others, 1993), and even fewer for the West Antarctic sea ice, including the pack ice in the Ross Sca. Consequently, an attempt was made to measure the albedo during a summer cruise on board the U.S. research vessel Nalhaniel B. Palmer in the Ross Sca in January and February 1999. Other parameters such as snow surface temperature, snow depth and snow grainsize were also routinely measured. In addition, one or two snow pits were dug at each of the ice-floe stations, and precise measurements of snow depth, stratigraphy, density, grain-size and morphology, salinity and temperature were recorded (Morris and Jeffrics, 2001).

## STUDY AREA, INSTRUMENTATION AND METHODS

## Study area

The cruise route is shown in Figure 1. The sea-ice cover in the Ross Sea is primarily first-year pack ice. Sea ice and its snow cover were used as the primary platform for albedo measurements. Temporal variation of the all-wave albedo and spectral albedo was measured from the northern edge to the southern edge of the pack ice along three lines of longitude: $165^{\circ} \mathrm{W}, 150^{\circ} \mathrm{W}$ and $135^{\circ} \mathrm{W}$. Albedo measurements were carried out on 24 first-year ice-floe stations and one land-fast ice site (station 9).

## Instrumentation and methods

The all-wave albedo and spectral albedo data discussed in this paper include data for both cloudy and clear-sky conditions. Almost all the measurements were carried out under cloudy sky conditions: only two sets of albedo data were obtained under very clear-sky conditions (stations 5 and 9). Downwelling solar irradiance was measured together with the all-wave albedo. Most measurements of all-wave albedo and spectral albedo were carried out from 1000 to 1600 h local solar time when the solar zenith angle was smallest. This choice was optimal for the measured results since the uncertainties of both measurement and model calculation are minimized (Schwander, and others, 1999). As the all-wave albedo generally varies considerably with time, depending on the cloud conditions (see next subsection), the two computers used for spectral albedo and all-wave albedo measurements were synchronized daily with the well-calibrated GMT clock on the ship. Datasets were time-stamped so that comparisons between the all-wave albedo retrieved from the spectral albedo and the all-wave albedo measured directly were possible.

## All-wave albedo

The total albedo (or all-wave albedo) was measured using two EPPLEY broadband, precision spectral pyranometers (PSPs) whose wavelength sensitivity was $200-4000 \mathrm{~nm}$. The two pyranometers were installed in the middle of a slender 2.2 m long aluminum bar and suspended about 1.4 m above the snow surface. Ninety-nine per cent of the upwelling irradiance received by the downward-looking pyranometer was from a ground field of view (GFOV) of 20 m . The glass domes of the pyranometers were cleaned before each deployment. As the results from the pyranometers are sensitive to the level of the instruments, care


Fig. 1. The cruise track of $R / V$ Nathaniel B. Palmer and albedo station locations during a summer cruise in the Ross Sea (January and February 1999). See Table 1 for corresponding days of year.
was taken to keep both pyranometers level. Onc PSP measured the downwelling irradiance, and the other the upwelling reflected irradiance. The two pyranometers were simultancously scanned every second, and the data saved at time intervals ranging from one second to several minutes, depending on the deployment period, using a CR10X datalogger. The save-time interval (STI) was generally 2060 s during our measurements, although a 1 s interval was sometimes used to study temporal variations of the all-wave albedo with high temporal resolution. The data stored were the average of the scans during the STI.

Time series of downwelling solar irradiance $\left(F_{\downarrow}(t)\right)$ were derived from the upward-looking pyranometer by converting the direct readings of voltage to flux units ( $\mathrm{W} \mathrm{m}^{-2}$ ). Likewise, the upwelling irradiance $\left(F_{\uparrow}(t)\right.$ ) was obtained from the readings of the downward-looking pyranometer, i.e.

$$
\begin{equation*}
F_{\downarrow}(t)=\frac{V_{\mathrm{U}}(t)}{S_{\mathrm{U}}}, \quad F_{\uparrow}(t)=\frac{V_{\mathrm{D}}(t)}{S_{\mathrm{D}}} \tag{1}
\end{equation*}
$$

where $V_{\mathrm{D}}(t)$ and $V_{\mathrm{U}}(t)$ are the time series of the readings in units of millivolts ( mV ) from the downward-looking and upward-looking pyranometers, respectively, and $S_{\mathrm{D}}$ and $S_{\mathrm{U}}$ are the sensitivities in units of $\mathrm{mV} \mathrm{W}^{-1} \mathrm{~m}^{2}$ of the down-ward-looking and upward-looking pyranometers, respectively. Thus the time series of all-wave albedo $(\alpha(t))$ were obtained from

$$
\begin{equation*}
\alpha(t)=\frac{F_{\uparrow}(t)}{F_{\downarrow}(t)}=\frac{V_{\mathrm{D}}(t)}{V_{\mathrm{U}}(t)} \frac{S_{\mathrm{U}}}{S_{\mathrm{D}}} . \tag{2}
\end{equation*}
$$

## Spectral albedo

Spectral albedo in the visible/near-infrared (VNIR) region was measured using a high-spectral-resolution, fiber-optic spectroradiometer manufactured by Analytical Spectral Devices, Inc. (ASD). The sensor measures spectral reflectance and spectral radiance between 333.65 and 1064.9 nm in 512 channels, with a spectral resolution of 3 nm . The spectral albedo was measured using a flux sensor (or a remote cosine receptor (RCR) in the manufacturer's terminology) connected to the spectroradiometer by a
fiber-optic cable. The base of the RCR had two built-in bubbles for horizontal alignment of the receptor. The RCR was mounted at one end of a slim level bar (about 60 cm long), with the other end attached to a tripod so that the receptor was about 40 cm from the snow surface. However, for most of the measurements of upwelling irradiance, the bar was held by hand with the RCR about 30 cm above the snow surfacc. Care was taken to insure that the RCR remained level during measurements of downwelling irradiance or upwelling irradiance. The spectral noise values - the electronic noise level of each of the 512 silicon photodiode detectors of the spectroradiometer - were collected before the sensors were exposed to any illumination, and were automatically subtracted from the data during the measurement. The spectral albedo was acquired by measuring the downwelling irradiance, followed immediately by measuring the upwelling irradiance using the RCR. Datasets were recorded by a mini-computer that was attached to the spectroradiometer. 'oo reduce system errors and errors caused by the fluctuations of the radiation field, several samples (each being the average of 10 scans) were collected for both upwelling and downwelling spectral irradiances.

For a Lambert surface, $99 \%$ of the upwelling irradiance that the cosine sensor receives comes from angles between $0^{\circ}$ (nadir) and $82^{\circ}$, i.e. within a GFOV of $2.135-2.846 \mathrm{~m}$ radius for a height of $30-40 \mathrm{~cm}$ above the surface. As the RCR optic and level bar are small and the level bar was always pointed towards the solar azimuth under sunny conditions, shadowing effects should be negligible. For a height of $3040 \mathrm{~cm}, 89.4^{-}$ $83.2 \%$ of the upwelling irradiance the RCR received was from the central part of the GFOV, with minimal shadowing effect from personnel. The shadowing effect duc to personnel is cstimated to be $1.1-1.8 \%$ under overcast sky conditions.

## Snow surface temperature

As the top 5 cm of the snow largely determines the albedo (Grenfell and others, 1994, the skin temperature on the snow surface was measured using a non-destructive method, an Exergen Model D-501-RS Infrared Microscamner (Li and others, 1999). To eliminate the effects of sky temperature and the surface emissivity on the temperature measurement, the head of the scanner was placed very ncar, but not touching, the surface to protect the sensor. This formed a black-body cavity between the snow surface and the surface of the microscanner, so that the measured temperature is the physical temperature of the skin of the snow surface. The microscanner is precise ( $\pm 0.1^{\circ} \mathrm{C}$ ) and quick in acquiring skin-temperature data. Skin-temperature data discussed in this paper are exclusively from the microscanner.

## CHARACTERISTICS OF SNOW SURFACE ON THE SEA ICE

Pack ice in the Ross Sea is covered by snow all year around. Even in summer, new snow was obscrved to fall on top of the old snow, although it metamorphosed very quickly under solar irradiation. Icy layers, melt clusters and ice lenses were ubiquitous in the snow cover due to melt and refreeze cycles. During the 1999 summer cruise, 25 ice floes were extensively examined and more ice floes were checked visually from the ship's bridge during the hourly ice observations: no melt ponds were observed. The conspicuous rotten edges of

Table 1. All-zeave albedo in eastern Ross Sea (albedo data in parentheses are for clear shy)

| Leg | Station No. <br> (Dav Wo.) | Position | Average | All-wave albedo Min. | Max. |
| :---: | :---: | :---: | :---: | :---: | :---: |
| first$\sim 165^{\circ} \mathrm{W}$ | 1 4) | $68^{\circ} \mathrm{S}, 165^{\circ} \mathrm{W}$ | 0.700 | 0.692 | 0.721 |
|  | 2 (5) | $69^{\circ} \mathrm{S}, 165^{\circ} \mathrm{W}$ | 0.769 | 0.741 | 0.808 |
|  | 3 (6) | $70^{\circ} \mathrm{S} 165^{\circ} \mathrm{W}$ | 0.754 | 0.729 | 0.771 |
|  | 47 | $71^{\circ} \mathrm{S}, 165^{\circ} \mathrm{W}$ | 0.788 | 0.752 | 0.825 |
|  | 5 (8) | $72{ }^{\circ} \mathrm{S}, 165.1^{\circ}$ W 0.79900788 |  | 0.78140.788) | 0.82000 .789 |
|  | 6 (9) | $\begin{gathered} 73 \mathrm{~S}, 1654^{\circ} \mathrm{W} \\ 73.7^{\circ} \mathrm{S} \\ 165.5^{\circ} \mathrm{W} \end{gathered}$ | 0.774 | 0.747 | 0.803 |
|  | 7 (10) |  | 0.832 | 0.819 | 0.855 |
|  | 8 (11) | $\begin{gathered} 74.8^{\circ} \mathrm{S} \\ 164.4^{\circ} \mathrm{W} \end{gathered}$ | 0.740 | 0.706 | 0.781 |
| Second$\sim 150^{\circ} \mathrm{W}$ | 9 (13) | $76^{\circ} \mathrm{S}, 151.3 \mathrm{w}$ | (0.827) | 0.814 | 0.830 |
|  | 10 (14; | $74^{\circ} \mathrm{S}, 150^{\circ} \mathrm{W}$ | 0.838 | 0.760 | 0.886 |
|  | 11 (15) | $\begin{gathered} 728^{\circ} \mathrm{S} \\ 1504^{\circ} \mathrm{W} \end{gathered}$ | 0.816 | 0.78 .5 | 0.855 |
|  | 12 (16) | $72{ }^{\circ} \mathrm{S}, 150{ }^{\circ} \mathrm{W}$ | 0.822 | 0.795 | 0.861 |
|  | 13 (17) | $71{ }^{\circ} \mathrm{S}, 150^{\circ} \mathrm{W}$ | 0.845 | 0.829 | 0.863 |
|  | 1418 | $70^{\circ} \mathrm{S}, 150 \mathrm{~W}$ | 0.842 | 0.785 | 0.871 |
|  | 15 (19) | $69^{\circ} \mathrm{S} .150^{\circ} \mathrm{W}$ | 0.778 | 0.765 | 0.799 |
|  | $16: 20$ ) | $68^{\circ} \mathrm{S}, 150^{\circ} \mathrm{W}$ | 0.766 | 0.725 | 0.802 |
|  | 17 (21) | $67^{\circ} \mathrm{S}, 150^{\circ} \mathrm{W}$ | 0.731 | 0681 | 0.777 |
|  | 18 (22) | $66^{\circ} \mathrm{S}, 149.9{ }^{\circ} \mathrm{W}$ | 0.797 | 0.774 | 0.837 |
| Third$\left(\sim 135^{\circ} \mathrm{W}\right)$ | 19 (26) | $70^{\circ} \mathrm{S}, 135^{\circ} \mathrm{W}$ | 0.763 | 0.747 | 0.779 |
|  | 20: 27 | $71^{\circ} \mathrm{S}, 135.1^{\circ} \mathrm{W}$ | 0.751 | 0.735 | 0.787 |
|  | 21 (28) | $72^{\circ} \mathrm{S}, 135^{\circ} \mathrm{W}$ | 0813 | 0.793 | 0.830 |
|  | 22 (29) | $73^{\circ} \mathrm{S}, 135^{\circ} \mathrm{W}$ | 0.857 | 0.846 | 0.871 |
|  | 23 30; | $\begin{gathered} 73.6^{\circ} \mathrm{S} \\ 134.9^{\circ} \mathrm{W} \end{gathered}$ | 0.821 | 0.772 | 0.844 |
|  | 24 (31) | $\begin{gathered} 74^{\circ} \mathrm{S}, 140^{\circ} \mathrm{W} \\ 75^{\circ} \mathrm{S}, 14.533^{\circ} \mathrm{W} \end{gathered}$ | 0.847 | 0.812 | 0.867 |
|  | 25 (33) |  | 0.791 | 0.775 | 0.820 |

almost all of the ice floes indicate that the decay of the sca ice occurred at the bottom and lateral cdges of the floes. Temporary wetting of the top snow layer was due to short periods of surface snow melting by solar heat, or by sea water splashed up onto the ice floe surface.

The snow surface temperature varied between $-0.3^{\circ}$ and $-3.8^{\circ} \mathrm{C}$, which is consistent with our observation that no slush or melt ponds werc obscrved on the surface of any floe. The difference in surface temperature on a single ice floe was as much as $1.9^{\circ} \mathrm{C}$ on station 11 January), and as little as $0.5^{\circ} \mathrm{C}$ on station 7 ( 10 January). Snow surface temperatures seemed to be primarily affected by the air temperature Li and others, 1999, with no observed correlation between snow depth and snow surface temperature.

## RESULTS OFALL-WAVE ALBEDO MEASUREMENTS

## Spatial variation

Daily-averaged albedos at all the stations are shown in Figure 2a. The average skin temperature of the snow surfacc on each floe is shown in Figure 2b. The threc scctions enclosed by dashed lines correspond to the threc legs of the cruise (Fig. l). The stations at the ends of each leg (stations 1,$8 ; 9,18 ; 19,25$ ) were at the margins of the pack ice. Albedos at the margins of the pack ice were lower than those in the central pack ice.

The daily-averaged albedos, the Julian dates and the positions at which they were acquired together with the variation range are summarized in Table 1. Because clouds deplete the longer wavelengths, the spectral distribution of downwelling irradiance is altered such that snow albedo is


Fig. 2. Daily-averaged all-wave albedo (a) and areaaveraged snow surface temperature (b) is time and latitude.
generally greater under overcast than under clear sky (Grenfell and Maykut, 1977). Variations in cloud conditions can also affect overcast albedos. From the albedo values shown in Table 1, we can sce that daily-averaged albedos were 0.7000 .857 , depending on the snow surface conditions and cloud conditions. The highest value was obtained at station 22 (29 January) on an ice floe where there was a surface layer of about 0.6 cm of recent snow, with an average grain diameter of about 0.2 mm . The lowest value was obtained at station 1 (4 January) on a relatively small ice floe (about 20 m in diameter) at the northern edge of the pack ice. The range in albedo for the pack ice in this region was similar to that obtained from the first-year sea ice in East Antarctica (0.65-0.85) (Allison and others, 1993).

The relationships between latitude and albedo and between latitude and snow surface temperature of ice flocs in the central pack under cloudy skies are shown in Figure 3: data from the most southerly and northerly ice stations (stations 1,$8 ; 9,18 ; 19,25$ ) along each of the three legs, and data taken under clear-sky conditions (station 10) are excluded. Mcasurements at station 5 included both clear and cloudy periods. The albedo shown in Figure 3 for this ice floc is the averaged value over the cloudy period of time. It can be seen that the albedo increases with latitude (Fig. 3a), while snow surface temperature decreases with latitude (Fig. 3b). The higher albedo at the more southerly stations is associated with the lower snow-surface temperatures and the resulting lower liquid-water content, together with smaller effective grainsize in the upper portion of the snow.

## Temporal change

A time series of irradiance and all-wave albedo measured at station 5 (7 8.January 1999) is shown in Figure 4. From midmorning local solar time ( $\sim 1000 \mathrm{~h})$ until almost noon, the sky was clear. The measured irradiance increased smoothly


Fig. 3. Pack-ice albedo and surface temperature values under cloudy-sky conditions: the linear regression shows the relationship between (a) average albedo and latitude and (b) average snowe surface temperalure and latilude. Stations at the pack-ice edge and on land-fast ice are excluded from this analysis.
with increasing solar elevation, but the albedo did not change significantly. At $23: 24$ GMT ( 1124 h , local time), it became cloudy, with the clouds moving rapidly over the ice floe while the snow surface conditions (topography, temperature, ctc.) remaincd almost the same. Clear and cloudy periods can be separated by the sudden variations in downwelling irradiance (lig. 4). Even before the cloud covered the site completely, the cloud scattered more light into the site, without blocking the direct solar beam. This caused the irradiance to increase slightly before it decreased abruptly as the cloud blocked direct sunlight to the site. When there were variations in the incident radiation due to changes in cloud conditions, the change in albedo was exactly opposite to that of downwelling irradiance, i.e. as


Fig. 4. Time series of downwelling irradiance and all-ziaze albedo measured on 7-8 January 1999 during clear- and cloudysky conditions. The station was located at $72^{\circ} 00.106^{\prime} S$, $165^{\circ} 03.242^{\prime}$ W. The snow depth was about 39 cm .
downwelling irradiance increased, albedo decreased. The average irradiance and albedo during the period of most stable clear-sky conditions (22:30 23:10 G.MT) and during the cloudy period are shown in Table 2, along with the irradiance fluctuation ranges during the 1 wo periods. From Table 2, we can see that the downwelling irradiance decreases by almost $23 \%$ due to the clouds. The increase of the average albedo due to clouds is about $1.4 \%$ relative to the average clear-sky albcdo. The largest increasc in albedo, which corresponded to the thickest cloud cover during the data acquisition, is $4.0 \%$. The increase of albedo due to clouds is a strong function of the cloud type, coverage and thickness (Grenfell and Maykut, 1977). Although these values are smaller than those obtained by Liljequist (1956) at Maudheim ( $5-10 \%$ ), Hanson (1960) at the South Pole ( $5-7 \%$ ), and Weller (1968) at Mawson ( $11 \%$ ), they are reasonablc, considering the broad temporal and spatial variability of the cloud coverage and thickness.

## EVALUATION OF MEASURED ALBEDO AND SPECTRAL ALBEDO

## Measured albedo compared to calculated all-wave albedo

To make sure that the measurements of solar radiation for the albedo calculation are reasonable, we compared the measured values with model calculations. The clear-sky measurements taken at ice station 9 (12-13 January) were used for this analysis. This station was the southernmost station of the cruise and also the only station on land-fast icc. At the begimning of the station occupation, the air temperature was $-5.4^{\circ} \mathrm{C}$ and the snow surface temperature was $-4.0{ }^{\circ} \mathrm{C}$. The downwelling irradiance and all-wave albedo are shown in Figure 5. The downwelling irradiance decreased continuously until the end of the measurements


12-13 January, 1999 (GMT)

Fig. 5. Time series of downwelling irradiance and all-ïave albedo measured on 12-13. Famuary during clear-sky conditions. Also shown are calculated irradiances with atmospheric attenuation ( $\tau=0.08$ ) and without almospheric attenuation $(\tau=0)$. The station was located at $76^{\circ} 01.287^{\prime}$ \&, $151^{\circ} 16.012^{\prime}$ W, with a multi-year snow cover about 154 cm thick. The excellent agreement between measured and calculated $(\tau=0.08)$ irradiance makes differences between the two curves almost indiscemible. Albedo remained relatively constant until the last third of the period, when it decreased stightlly from 0.828 to 0.814 .

Table 2. Downwelling irradiance and albedo for clear and cloudy periods, 7-8 January 1999

| Sky conditions | Iradiance |  | A/bedo |  |
| :---: | :---: | :---: | :---: | :---: |
|  | Average | Range | Average | Range |
|  | $W^{+} \mathrm{m}^{-2}$ | $\mathrm{Hm}^{2}$ |  |  |
| Clcar | 75.1 | $741 \sim 770$ | 0.788 | 0.788~0.789 |
| Ciloudy | 580 | $353 \sim 896$ | 0.799 | 0.784~10.820 |

as the solar elevation angle decreased after local noon. However, the albedo did not change significantly. During the last third of the measurement period, the albedo decreased slightly.

Because the aumosphere was clear, we used a simple solar-zenith-angle model to calculate irradiance values for this time period to test the correctness of the irradiance measurements. Supposing at starting time $t_{0}$ the theoretical value of downwelling irradiance $F_{1}^{*}\left(t_{0}\right)$ is exactly the same as the measured value $F_{( }\left(t_{0}\right)$, then for any time $t$ later, the theoretical irradiance valuc is

$$
\begin{align*}
F_{\downarrow}^{\prime}(t) & =\frac{\cos [\theta(t)]}{\cos \left[\theta\left(t_{0}\right)\right]} F_{1}^{\prime}\left(t_{0}\right) \exp \left\{-\tau\left[\sec ^{-1} \theta(t)-\sec ^{-1} \theta\left(t_{0}\right)\right]\right\} \\
& =\frac{\cos [\theta(t)]}{\cos \left[\theta\left(t_{0}\right)\right]} F_{v}\left(t_{0}\right) \exp \left\{-\tau\left[\sec ^{1} \theta(t)-\sec ^{-1} \theta\left(t_{0}\right)\right]\right\} \tag{3}
\end{align*}
$$

where $\theta(t)$ is the solar zenith angle at time $t$, and $\tau$ is the apparent bulk optical depth of the clear-sky atmosphere. Based on the results of Wendler 1986 and Li and others (1987), $\tau$ was chosen to be 0.08 . The solar-\%enith-angle model devised by Woolf (1968) was used to calculate the temporal variation of the solar zenith angle. The results of the calculated surface solar irradiance with and without atmospheric attenuation are compared with actual measurements in Figure 5. For the case of no atmospheric attenuation ( $\tau=0$ ), the maximum deviation of calculated from measured surface solar irradiance is $1.43 \%$ and occurs at 1:27:40 GM'T on 13 January when the measured value is $548 \mathrm{~W} \mathrm{~m}^{2}$ and the calculated value is $556 \mathrm{~W} \mathrm{~m}^{-2}$. For the case $\tau=0.08$, the maximum deviation of calculated from measured surface solar irradiance is $0.26 \%$ and occurs at 1:22:00 GMT on 13 . January when the measured value is $5.54 \mathrm{~W} \mathrm{~m}^{-2}$ and the calculated valuc is $552 \mathrm{~W} \mathrm{~m}^{-2}$. This comparison demonstrates that even using a simple approximation there is good agreement between measured and calculated values, especially when the atmospheric attenuation is considered. Neverthcless, an effort is being made to model albedo using a more complicated radiative transfer method.

## Comparison of all-wave albedo and spectral albedo measurements

Spectral albedo ( $\alpha_{\lambda}$ ) was measured simultaneously with all-wave albedo. In order to determine how well these two sets of measurements agree, all-wave albedos derived from spectral albedo data were compared to the measured values. First, the spectrally averaged visible and near-infrared
(333.65 1064.9 nm ) VNIR albedo ( $\alpha_{\text {vnir }}$ ) was obtained from

$$
\begin{equation*}
\alpha_{\mathrm{vnir}}=\frac{\int_{333.65}^{1064.9} \alpha_{\lambda} F_{\lambda}^{\downarrow} \mathrm{d} \lambda}{\int_{333.65}^{1064.9} F_{\lambda}^{\downarrow} \mathrm{d} \lambda}=\frac{\sum_{i=1}^{512} \alpha_{i} F_{i}^{\downarrow} \Delta \lambda_{i}}{\sum_{i=1}^{512} F_{i}^{\downarrow} \Delta \lambda_{i}} \tag{4}
\end{equation*}
$$

where $\alpha_{i}, F_{i}^{\downarrow}$ and $\Delta \lambda_{i}$ are the albedo, downwelling irradiance and bandwidth of the $i$ th channcl of the 512 chanmels of the spectroradiometer. The all-wave albedo $\alpha_{\text {int }}$ can be estimated through intcgration over the measured VNIR spectral albedo with extrapolation to the unmeasurcd spectral regions for clear-sky conditions:

$$
\begin{align*}
\alpha_{\text {int }} & =\left\{\int_{280}^{334} F_{\lambda}^{\downarrow} \alpha_{334} \mathrm{~d} \lambda+\int_{334}^{10660} F_{\lambda}^{\downarrow} \alpha_{\lambda} \mathrm{d} \lambda\right.  \tag{5}\\
& \left.+\int_{1060}^{4000} F_{\lambda} \frac{\left(\alpha_{1060}+0.07\right)}{2} \mathrm{~d} \lambda\right\} / \int_{280}^{4000} F_{\lambda} \mathrm{d} \lambda
\end{align*}
$$

where $\quad \alpha_{\lambda}=\alpha_{334}$ is assumed for $\lambda<334 \mathrm{~nm}$ and $\alpha_{\lambda}=\left(\alpha_{1060}+0.07\right) / 2$ is assumed for $\lambda>1060 \mathrm{~nm}$ (Allison and others, 1993). Other values of $\alpha_{\lambda}$ within $\lambda \in[331,1060]$ are the measured spectral albedo, and $\alpha_{\lambda}=0.07$ is the albedo of open water. Equation (5) provides a method for deriving $\alpha_{\text {int }}$ from the measured spectral albedo when the proportions of the downwelling irradiance in the ultraviolet (UV), VNIR and shortwave infrared (SWIR) spectral regions are known. The proportions of downwelling clearsky spectral irradiance in different wavelength regions were calculated using the radiative transfer algorithm of Stamnes and others (1988).

To derive the all-wave albedo from the measured spectral albedo for cloudy conditions, the proportion of the downwelling SWIR irradiance was calculated by subtracting the LV and VNIR irradiance from the all-wave irradiance. Compared with near-infrared radiation, the effect of clouds on the UV radiation is negligible, so the spectral irradiance $F_{\lambda}^{\downarrow}$ in the UV wavelength region $280-334 \mathrm{~nm}$ was taken to be the corresponding clear-sky downwelling spectral irradiance calculated using the radiative transfer algorithm (Stamnes and others, 1988). Within the VNIR wavelength region, we used the measurements from the $A S D$ spectroradiometer. The all-wave downwelling solar flux was measured by the upward-looking pyranometer (PSP). To account for the difference in the sensitivities between the PSP pyranometer and the ASD spectroradiometer, we defined instrument-dependent parameters $\kappa_{1}$ and $\kappa_{2}$ to relate the measured downwelling solar irradiance $F^{\mathrm{PSP}}$ and spectral irradiance $F_{\lambda, \text { ASD }}^{\downarrow}$ with the theoretical spectral irradiance $F_{\lambda}$, derived from the radiative transfer model for clear-sky conditions:

$$
\begin{equation*}
\kappa_{1}=\frac{F^{\mathrm{PSP}}}{\int_{280}^{4000} F_{\lambda}^{\downarrow} \mathrm{d} \lambda}, \quad \kappa_{2}=\frac{\int_{334}^{1060} F_{\lambda, \lambda \mathrm{SD}}^{\downarrow} \mathrm{d} \lambda}{\int_{334}^{1060} F_{\lambda}^{\downarrow} \mathrm{d} \lambda} \tag{6}
\end{equation*}
$$

For the ideal case, $\kappa_{1}=\kappa_{2}=1$. For the pyranometer and the spectroradiometer, we used $\kappa_{1}=1.128$ and $\kappa_{2}=0.883$, as obtained using a clear-sky casc at station 5 (8 January).


Fig. 6. (a) Spectral albedos measured under clear-sky conditions at station 5 ( 8 January) and station 9 (13. January). (b) Spectral albedos measured under cloudy conditions. al station 10 (14 January) and station 12 ( 16 . January ).

Then the equivalent of Equation (3) takes the form for the gencral cascs,

$$
\begin{aligned}
\alpha_{\mathrm{int}}= & \kappa_{1} \alpha_{331} \int_{280}^{334} \frac{F_{\lambda}^{\downarrow}}{F^{\mathrm{PSP}}} \mathrm{~d} \lambda \\
& +\frac{\kappa_{1}}{\kappa_{2}} \int_{334}^{1060} \frac{F_{\lambda, \mathrm{ASD}}^{\downarrow} \alpha_{\lambda}}{F^{\mathrm{PSP}}} \mathrm{~d} \lambda+\frac{\left(\alpha_{1060}+0.07\right)}{2} \\
& \left(1-\kappa_{1} \int_{280}^{334} \frac{F_{\lambda}^{\downarrow}}{F^{\mathrm{PSP}}} \mathrm{~d} \lambda=\frac{\kappa_{1}}{\kappa_{1}} \int_{3,334}^{1060} \frac{F_{\lambda, \mathrm{ASD}}^{\downarrow}}{\left.F^{\mathrm{PSP}} \mathrm{~d} \lambda\right) .}\right.
\end{aligned}
$$

Spectral albedos obtained under clear-sky conditions at station 5 (8 January) on pack ice and at station 9 ( 13 January) on land-fast ice are shown in Figure 6a. Spectral albedos obtained under cloudy sky conditions at station 10 (14 January) on central pack icc and at station 12 (16 January) are shown in Figure 6b. The snow characteristics, solar zenith angle and albedo values are summarized in Table 3.

At station 5, the snow surface was made up of melt clusters ranging from 0.3 to 0.6 mm , with an average grain-size $(D)$ of 0.4 mm and snow density $\rho=320 \mathrm{Mg} \mathrm{m}{ }^{3}$ for the 6 cm thick surface layer. At station 9, the snow surface was wind-packed and icy and made up of 0.1 mm snow grains. No snow-density data are available at this station. Although the sky was clear when the spectral albedo was measured at stations 5 and 9, it was much clearer at station 9 than at station 5 (see Figs 4 and 5. There was a very thin but stable layer of cirrus clouds at station 5 during the clear period in the day. From Figure 6a we can see that the spectral albedo was smaller at station 5 than at station 9 until the wavelength exceeded 994 nm . This

Table 3. Snow paramelers and refrieved albedos for selected sites where spectral albedo was measured

| Parameters | Clearsky |  | Cloudy shy |  |
| :---: | :---: | :---: | :---: | :---: |
|  | Stution 5 | Stationg | Station 10 | Station I? |
| Surface temperature ( C | -1.5 | $-4.0$ | -2.3 | -1.4 |
| Snow depth (cm) | 39 | 154 | 34 | 22 |
| Snow grain-size (mm) | 0.4 | 0.1 | 0.31 .0 | 0.21 .0 |
| Snow density (M gm ${ }^{3}$; | 0.32 |  | 0.33 | 0.39 |
| Solar zenith angle () | 50 | 55 | 5 | 50 |
| $\alpha_{\text {vuir }}$ | 0.847 | 0.892 | 0.900 | 0.820 |
| $\alpha_{\text {int }}$ | 0.773 | 0.797 | 0.838 | 0.786 |
| $\alpha$ | 0.788 | 0829 | 0.834 | 0.818 |
| $\Delta_{1}-\alpha_{\text {vnir }}-\alpha$ | 0.059 | 0063 | 0.066 | 0002 |
| $\Delta_{2}-\alpha_{\text {int }}-\alpha$ | $-0.015$ | -0.032 | 0.004 | 0.032 |
| $\Delta_{3}=\alpha_{\mathrm{vair}}-\alpha_{\mathrm{iat}}$ | 0.074 | 0.095 | 0.062 | 0.034 |

was mainly due to the smaller grain-size ( $D=0.1 \mathrm{~mm}$ ) at station 9 (Wiscombe and Warren, 1980; Dozier, 1989;.

The surface of the ice floc at station 10 was rough. Snow depth across the floc was horizontally very inhomogencous, ranging from 3 to 60 cm . The grain-size of the top 1 cm layer was $D=0.3 \mathrm{~mm}$, wilh coarse-grained snow ( $D=1 \mathrm{~mm}$ ) directly below. Average density for the surface layer $(6 \mathrm{~cm})$ was $330 \mathrm{Mg} \mathrm{m}{ }^{3}$. The snow surface of the ice floe was flater at station 12 than at station 10 . On the very top of the snow cover was a 0.5 cm thick layer of recent snow with $D=0.2 \mathrm{~mm}$, under which was a cernented icy layer of 3 cm with $D=0.6 \mathrm{~mm}$. Under the icy layer were clusters of rounded particles with $D=1 \mathrm{~mm}$. The averaged density of the upper 6 cm was $390 \mathrm{Mg} \mathrm{m}{ }^{3}$. The larger spectral albedo at station 10 (Fig. 6b) is mainly due to the finer grain-size ( $D=0.2 \mathrm{~mm}$ ) of the top layer.

Derived all-wave albedo values ( $\alpha_{\mathrm{int}}$ ) using Equations (5) and (7) for stations 5, 9, 10 and 12 are listed in Table 3. The differences between $\alpha_{\text {vnir }}$ and $\alpha_{\text {int }}$ are smaller (3.4-6.2\%) under cloudy than under clear-sky conditions (7.4 9.5\%) because the SWIR proportion was smaller under cloudy conditions. For example, at station 10 , the SWIR accounted for only $7.6 \%$ of the total downwelling irradiance, but at station 5 where the sky was clear, the SWIR accounted for $19.1 \%$ of the total downwelling radiation.

The difference between the measured all-wave albedo $\alpha$ and those derived from measured V.VIR spectral albedo is within $\pm 3.2 \%$. Since all-wave albedo and spectral albedo were measured on the same ice floes, yet at separate locations, these results seem reasonable given the uncertaintics associated with spectral extrapolation.

## CONCLUSIONS

The current investigation has focused on regional all-wave and spectral albedo measurements of the snow cover on the austral-summer sea ice in the Ross Sea. No obvious correlation between snow thickness and skin temperature of the snow surface emerged from our data. Snow surface temperature was primarily correlated with air temperature Li and others, 1999).

The measured all-wave downwelling irradiance agreed well with a simple approximation using solar zenith angle and bulk optical thickness for clear-sky conditions. Measured irradiance has not been compared with calculations for over-
cast conditions, duc to lack of quantitative cloud-cover data. Albedo varied greatly (0.70 0.86 between different places in the Ross Sea. This large variation in albedo was also obscrved in the first-year pack ice in East Antarctica (0.65-0.85). Clearand cloudy-sky albedos measured at the same site indicate that the average increase due to clouds is $1.4 \%$, with the largest increase being $4.0 \%$.

All-wave albedos were observed to be lower at the margin of the pack ice than in the central pack icc. Over the central pack ice, it was observed that the higher the latitude, the lower the snow surface temperature and the higher the all-wave albedo.

There was no major difference between spectral albedos under clear-sky and cloudy conditions. Both are mainly controlled by the snow grain-size, especially the snow grain-size of the top snow laycr. All-wave albedos calculated from direct measurements of all-wave irradiance agreed well with those derived from integration of visible and ncar-infrared spectral albedos which included extrapolation to UV and SWIR regions for both clear- and cloudy-sky conditions. This indicates that our all-wave and spectral albedo measurements are reliablc.

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